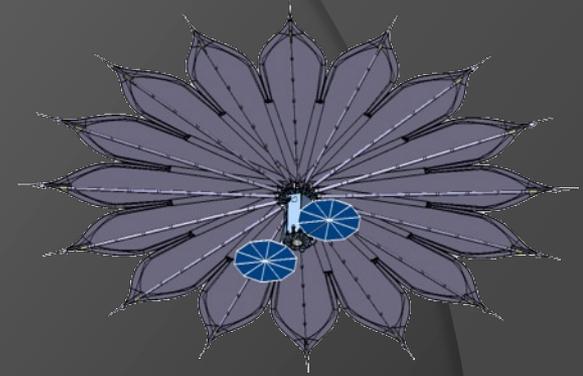


# EXOPLANET PROBE – STARSHADE

## STUDY UPDATE



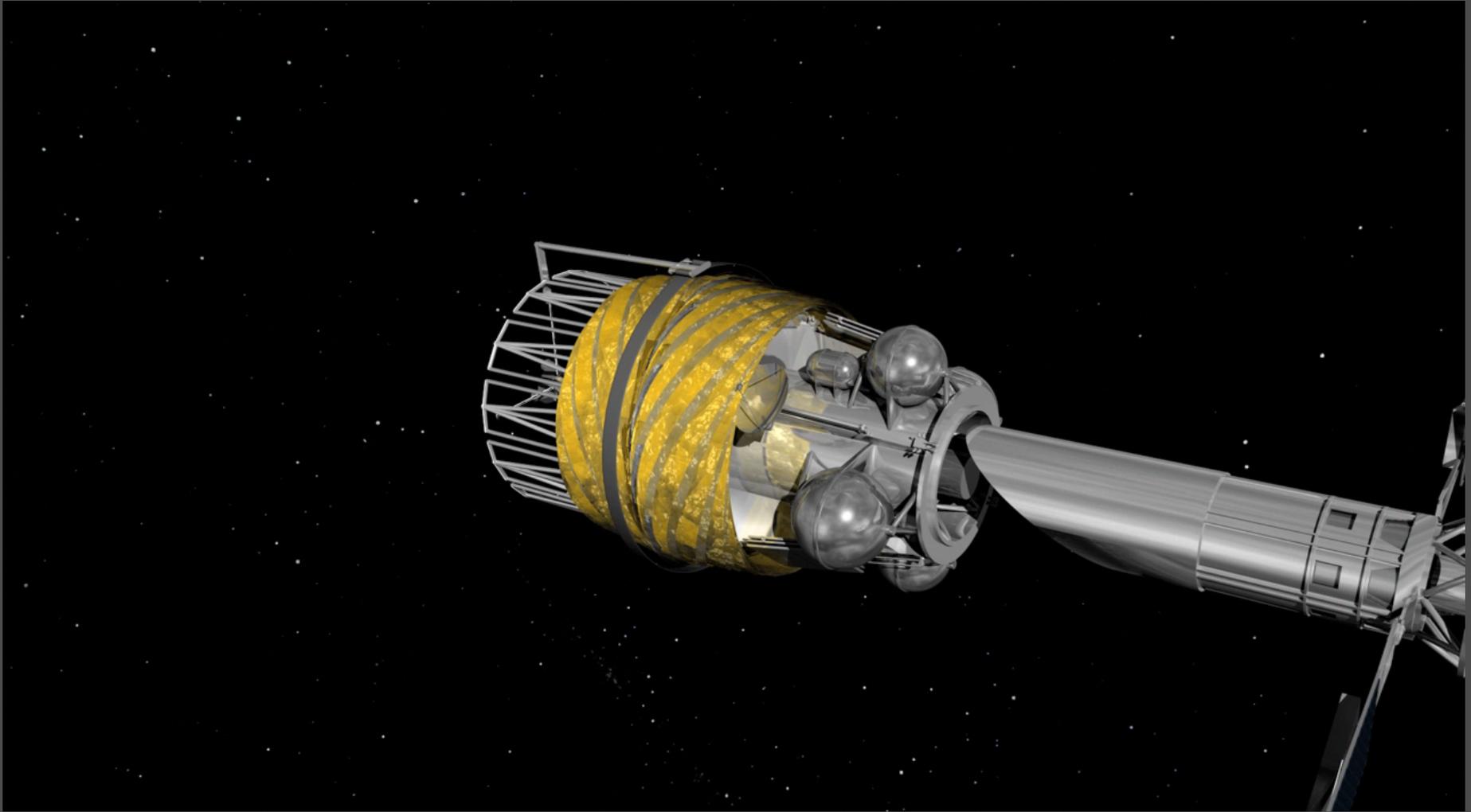
**Aki Roberge (NASA GSFC)**

on behalf of the Exoplanet Probe – Starshade STDT  
and Design Team

**STDT:** S. Seager (MIT), W. Cash (Colorado), S. Domagal-Goldman (NASA GSFC), N. J. Kasdin (Princeton), M. Kuchner (NASA GSFC), A. Roberge (NASA GSFC), S. Shaklan (NASA JPL), W. Sparks (STScI), M. Thomson (NASA JPL), M. Turnbull (GSI)

**Design Team:** D. Lisman, E. Cady, S. Martin, D. Webb, ... (NASA JPL)

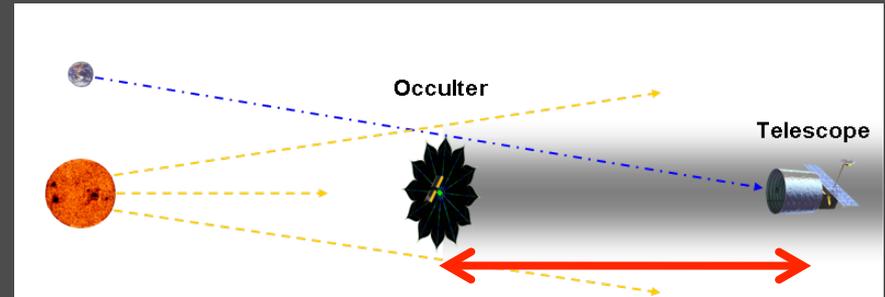
# Starshade primer



<http://planetquest.jpl.nasa.gov/video/15>

# Starshade strengths

- Contrast and inner working angle decoupled from telescope size
- Inner working angles can be changed by altering telescope/occulter separation
- No outer working angle
- 360 degree suppressed field of view



W. Cash (Colorado)

# Starshade strengths

- High quality telescope not required
  - Segments & obstructions not a problem
  - Wavefront correction not needed
- Broad bandpass, high total throughput
- No constraints on other astronomical instruments



NASA / STScI



NASA / Swift

# Starshade drawbacks

- ⦿ Full-scale end-to-end system test on the ground not possible
  - Sub-scale lab and field tests possible (more later)

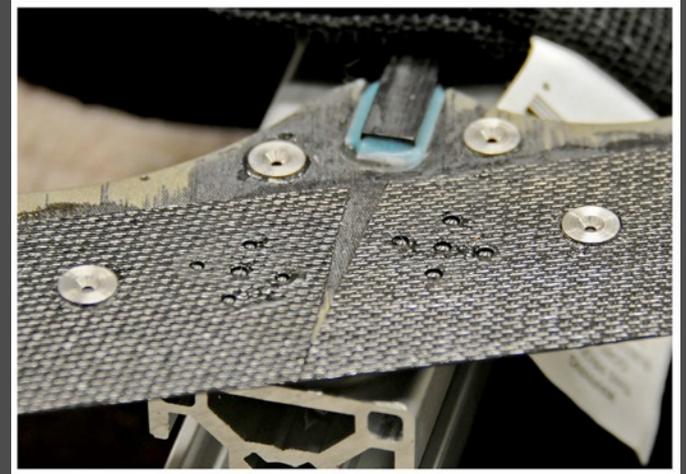


T. Glassman / NGAS

- ⦿ Long times between observations while moving starshade
  - But can do other astronomy in the intervals
- ⦿ Limited number of starshade movements

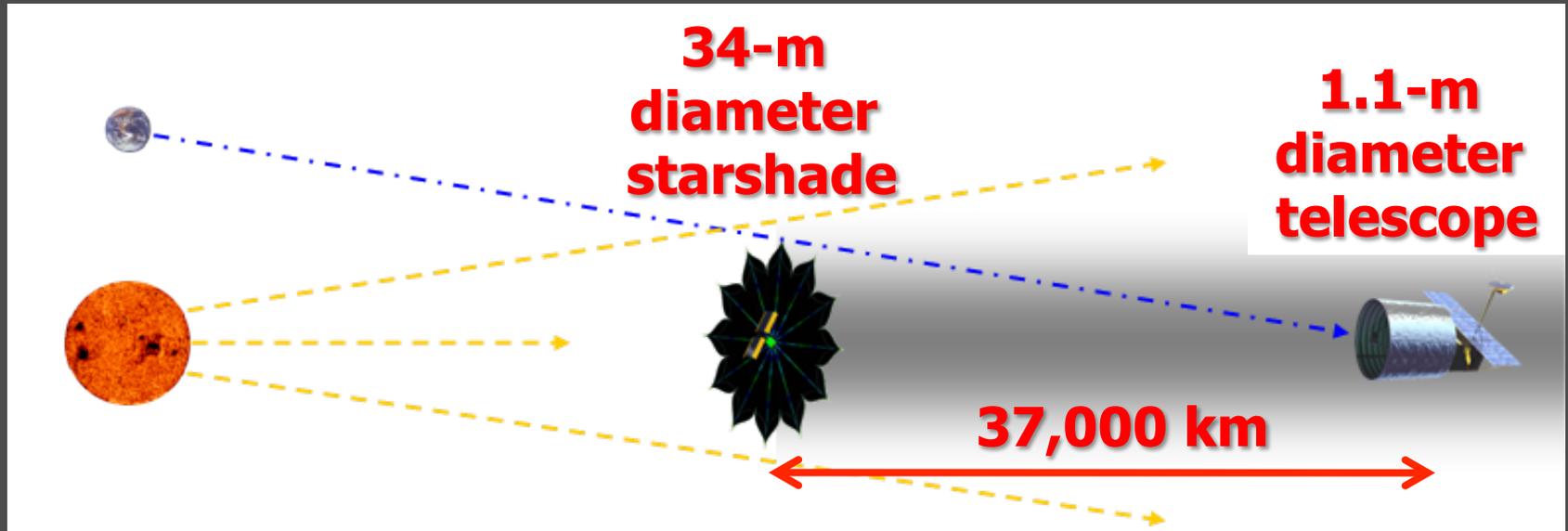
# Technical challenges

- Precise edge profile  
(~ 50  $\mu\text{m}$  tolerance) required  
over large structure
- On-orbit deployment of  
large structure
- Precise alignment between starshade and  
telescope needed ( $\pm 1$  meter tolerance)
- Cost models not fully developed



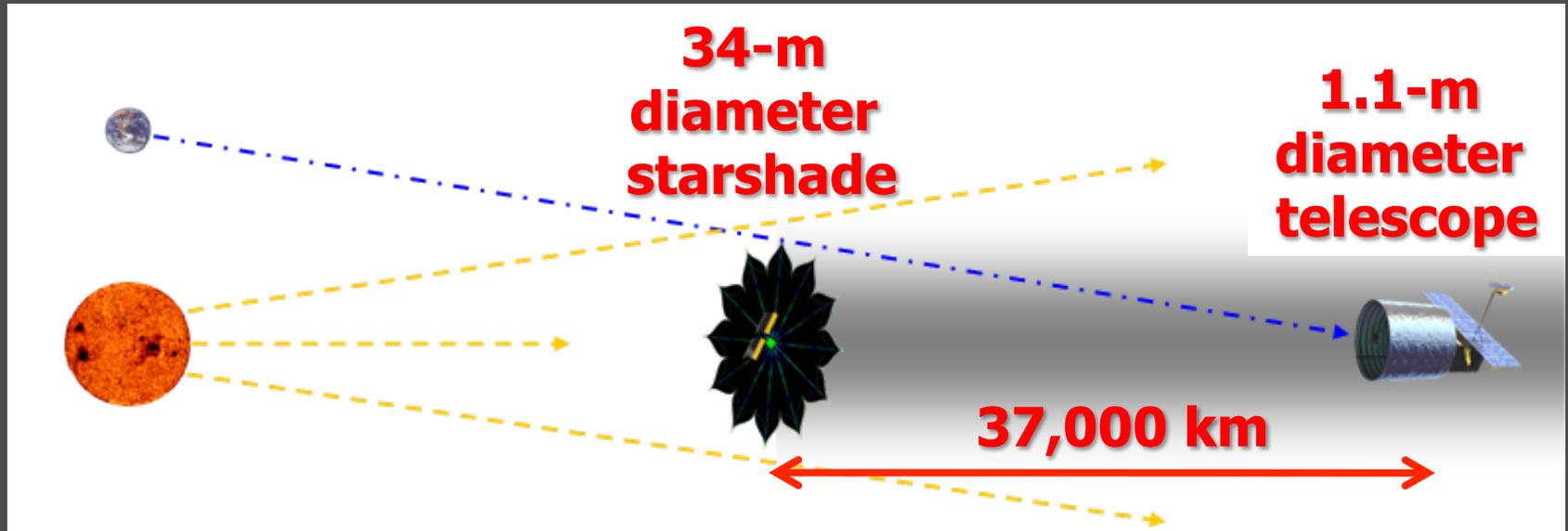
NASA / JPL / Princeton

# Probe baseline design specs



- ⦿ Off-the-shelf on-axis optical telescope
- ⦿ Earth-leading orbit
- ⦿ Imager and low resolution spectrograph
- ⦿ Slewing telescope, not starshade

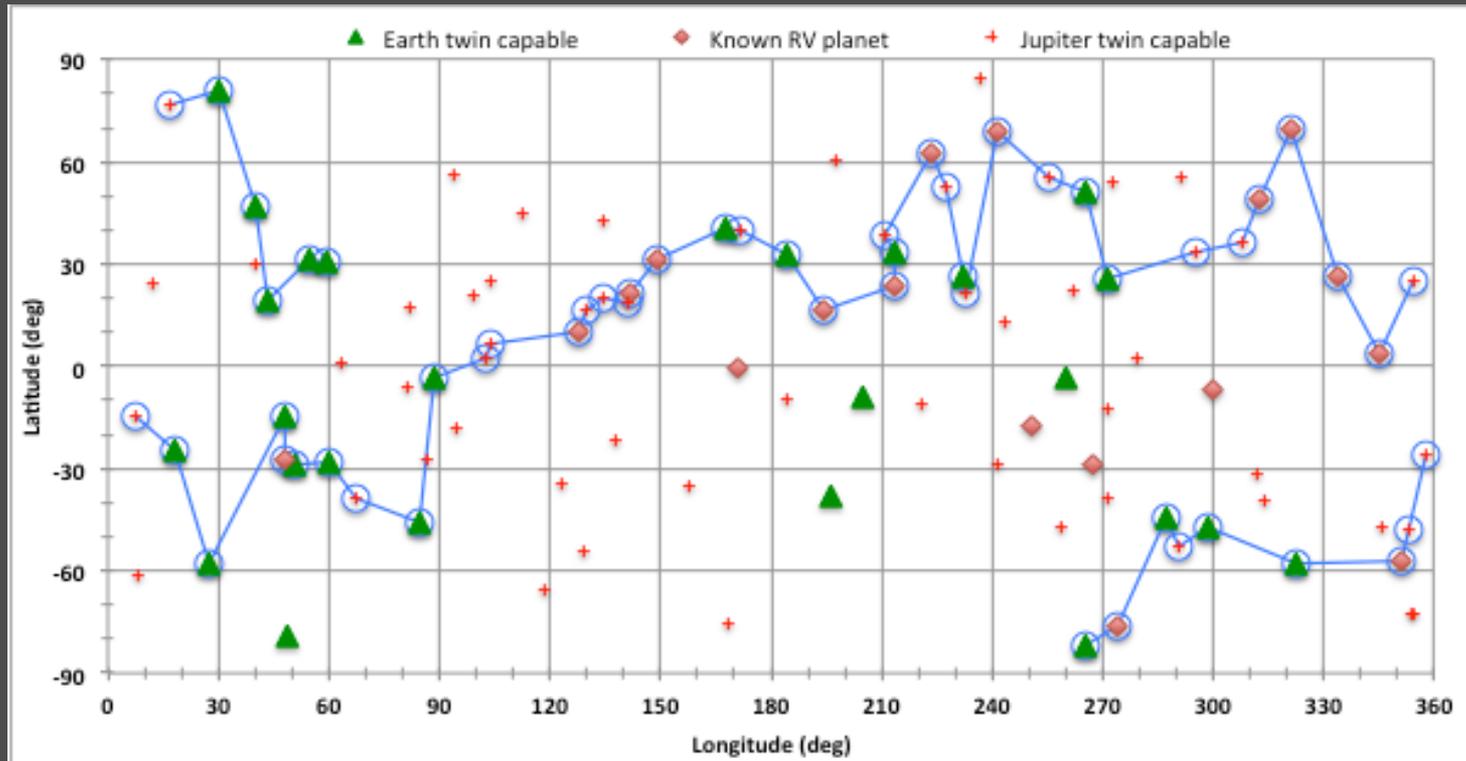
# Probe baseline design specs



- ◎ Primary operating mode
  - 500 – 850 nm bandpass
  - 95 milli-arcsecond inner working angle
  - Limiting fractional planet brightness  $\sim 9 \times 10^{-11}$
- ◎ Other bands with different IWAs for follow-up

# Preliminary observing strategy

- First 18 months in “reconnaissance mode”
  - Multi-color imaging only to find candidates
- Second 18 months for revisits and spectroscopy



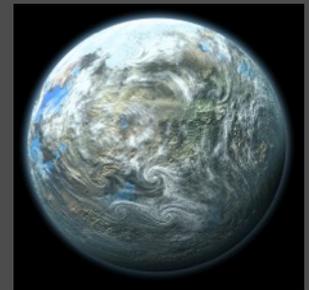
D. Lisman (NASA JPL)

# Preliminary science yield predictions

- ◎ In 18 months, observe 55 stars
  - Assuming 1 zodi of exozodi dust ...
  - Can detect Jupiter-twins around all stars
  - 14 stars with detectable known giant planets from radial velocity surveys
  - Possibility of detecting Earth-analog exoplanets around 22 stars
- ◎ Remainder of mission for revisits, follow-up spectroscopy, maybe disk observations



NASA

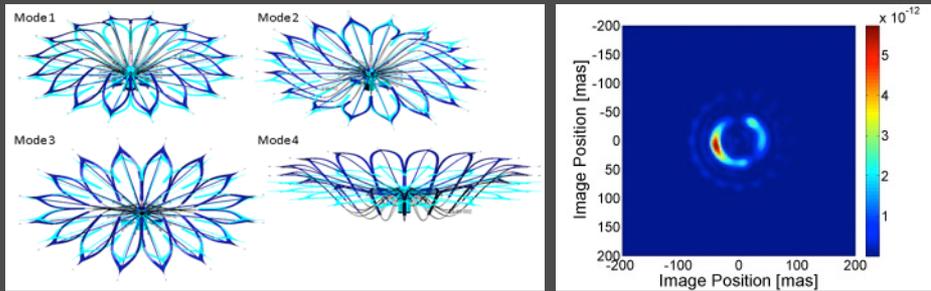


L. Cook

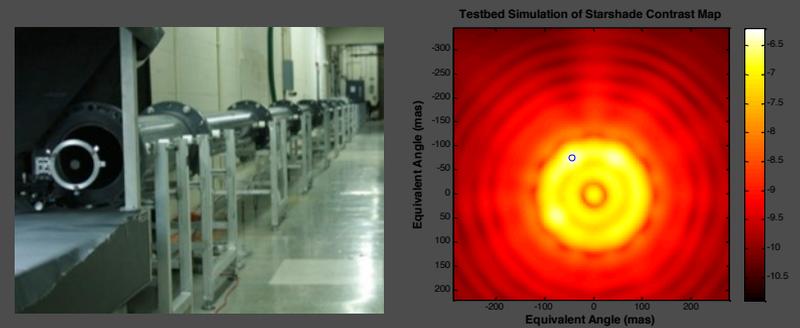
# Technology demonstrations

## Performance Modeling and Testing

### Optical models with distortions



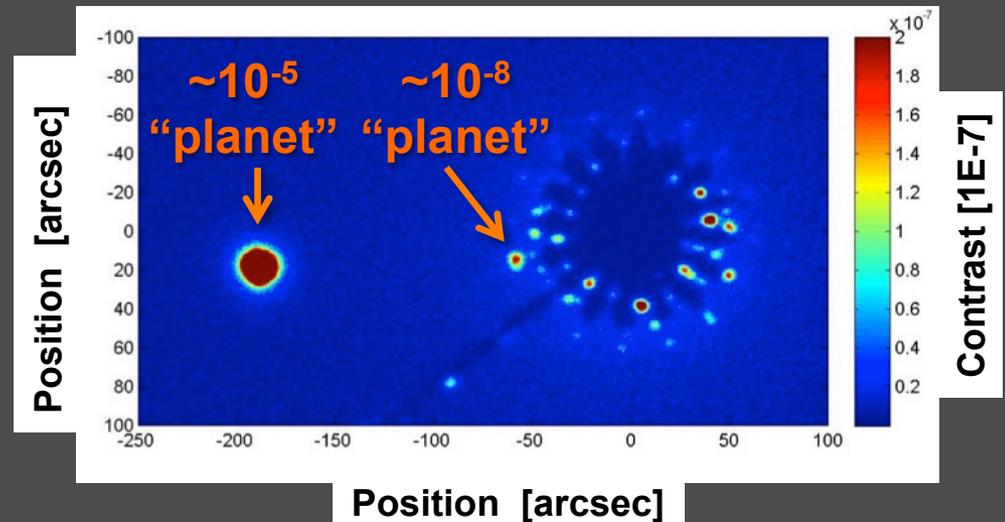
### 0.1% scale lab testing



### ~ 1% scale field testing



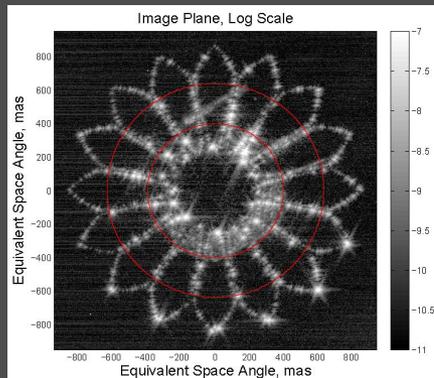
T. Glassman / NGAS



# Technology demonstrations

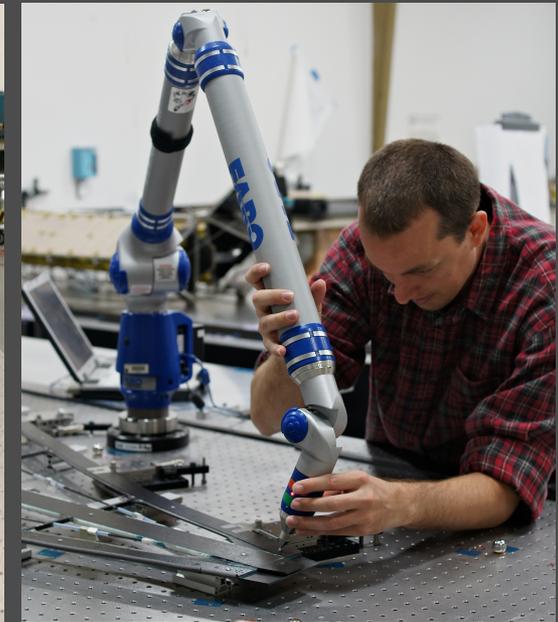
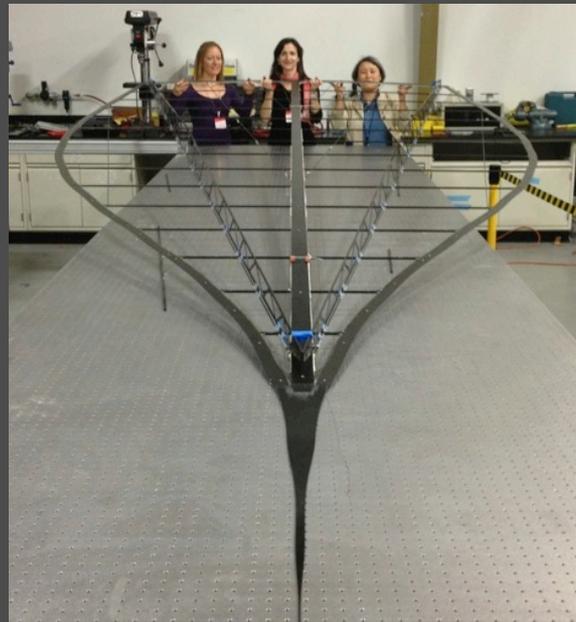
## Precision petal manufacturing

### Sub-scale full starshade



Sirbu, Kasdin, & Vanderbei 2013

### Full-scale petal with required edge profile

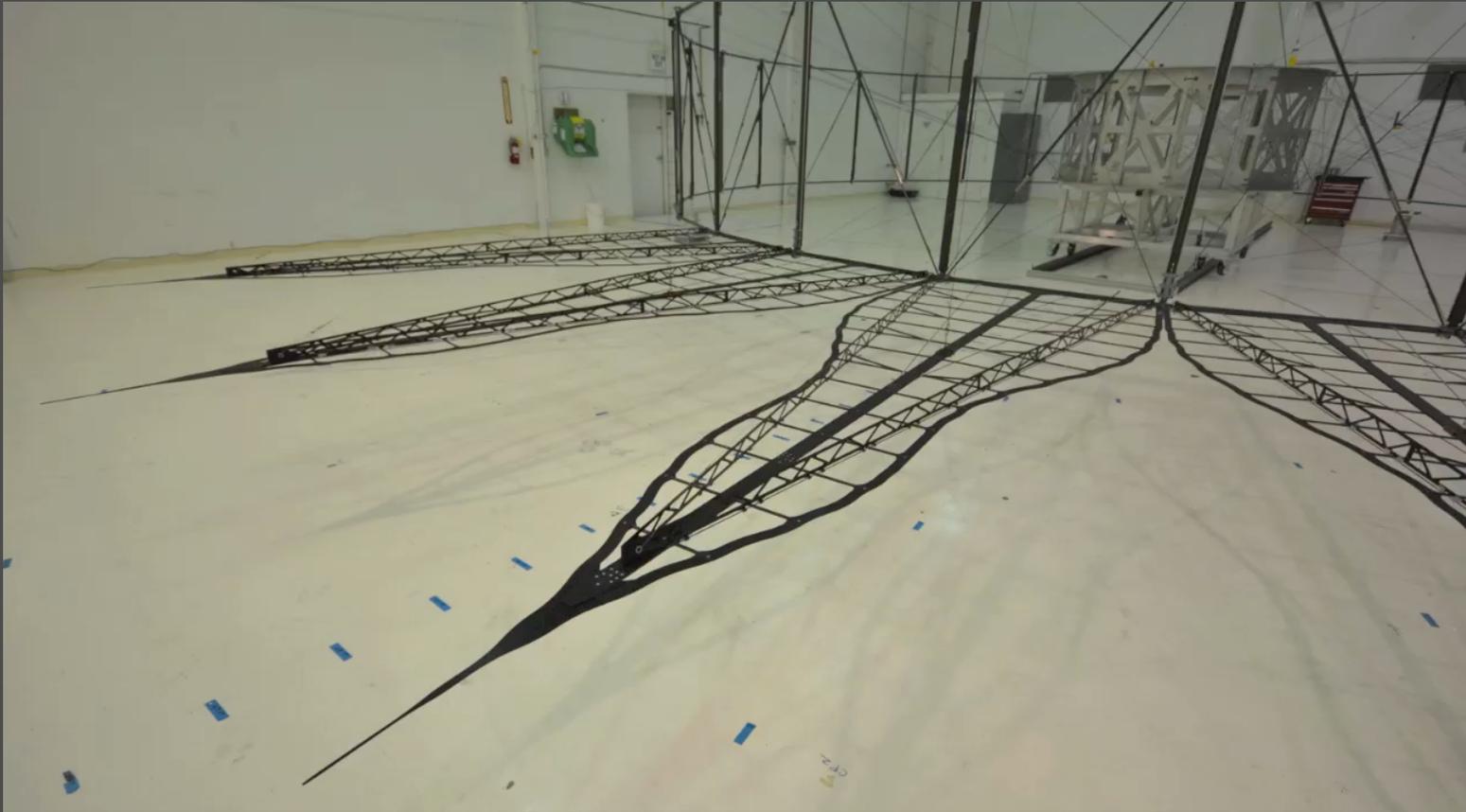


D. Lisman

- Development of knife-edge to control edge scatter underway

# Technology demonstrations

## Starshade stowage and deployment



<http://www.youtube.com/watch?v=G68sqqRhP2E>

# Next steps

- ⦿ Refine baseline mission design and science yield simulations
- ⦿ Continued technology development
  - Through competed NASA technology programs; some STDT members participating
- ⦿ Work on cost modeling with Aerospace Corp.
- ⦿ Consider general astrophysics science cases
- ⦿ Make design for starshade with NRO telescope